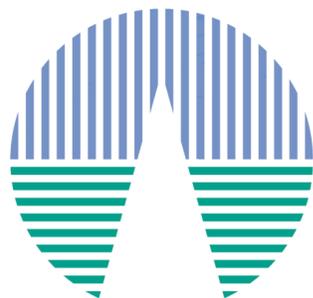


THE VIRTUAL OBSERVATORY (VO)

ALBA ALLER EGEA
CENTRO DE ASTROBIOLOGÍA (CAB, CSIC-INTA)

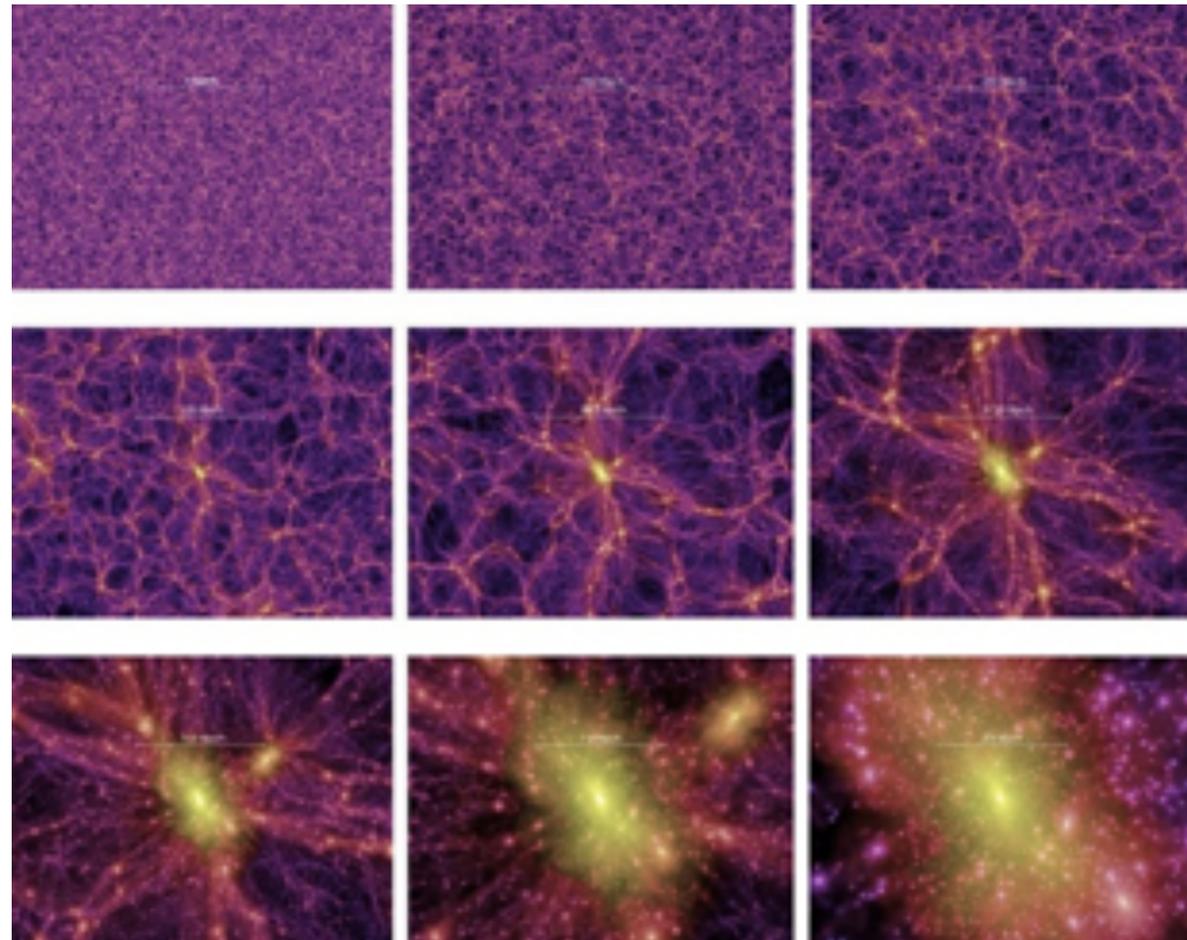


CENTRO DE ASTROBIOLOGÍA · CAB

ASOCIADO AL NASA ASTROBIOLOGY PROGRAM



DATA, DATA, AND MORE DATA!



ASTRONOMICAL ARCHIVES



- ESDC » Home
- Home
- About ESDC
- Archival Research Visitor Programme
- Newsletter
- Science Archives
- Archive Image Browser
- ESASky
- DOIs
- User Survey Results
- Videos
- Scientific Tutorials
- Publications
- VOSpec
- Euro-VO Registry
- Archives User Groups
- Contact Us

ESAC SCIENCE DATA CENTRE

ESDC Statistics

Monthly Users (*) 16 879	Monthly Downloaded (*) 101. TB	Archive Total Size 662.2 TB
---------------------------------	---------------------------------------	------------------------------------

* Monthly averages in 2021

Astronomy Science Archives

- cheops
- esasky
- exosat

European Southern Observatory

ESO — Reaching New Heights in Astronomy

Public Science User Portal Intranet Contact Site Map Search Go!

Science Archive Facility

Welcome to the ESO Science Archive Facility

The ESO Science Archive Facility contains data from ESO telescopes at La Silla Paranal Observatory, including the APEX submillimeter telescope on Llano de Chajnantor. In addition, the raw UKIDSS/WFCAM data obtained at the UK Infrared Telescope facility in Hawaii are available.

The Principal Investigators of successful proposals for time on ESO telescopes have exclusive access to their scientific data for the duration of a proprietary period, normally of one year, after which the data becomes available to the community at large. Please read the [ESO Data Access Policy](#) statement for more information, along with the [relevant FAQs](#).

Browsing the archive does not require authentication. Please [acknowledge the use of archive data](#) in any publication.

There are three main ways to access the archive, varying for content and presentation/interface: the usual Raw Data query form, the innovative Science Portal to browse and access

The Mikulski Archive for Space Telescopes is an astronomical data archive focused on the optical, ultraviolet, and near-infrared. MAST hosts data from over a dozen missions like Hubble, Kepler, TESS, and soon JWST.

- ### On This Page
- Missions
 - High Level Science Products
 - Search Tools
 - Catalogs
 - APIs
 - Publishing Resources
 - Latest MAST News

Missions



Hubble



Webb



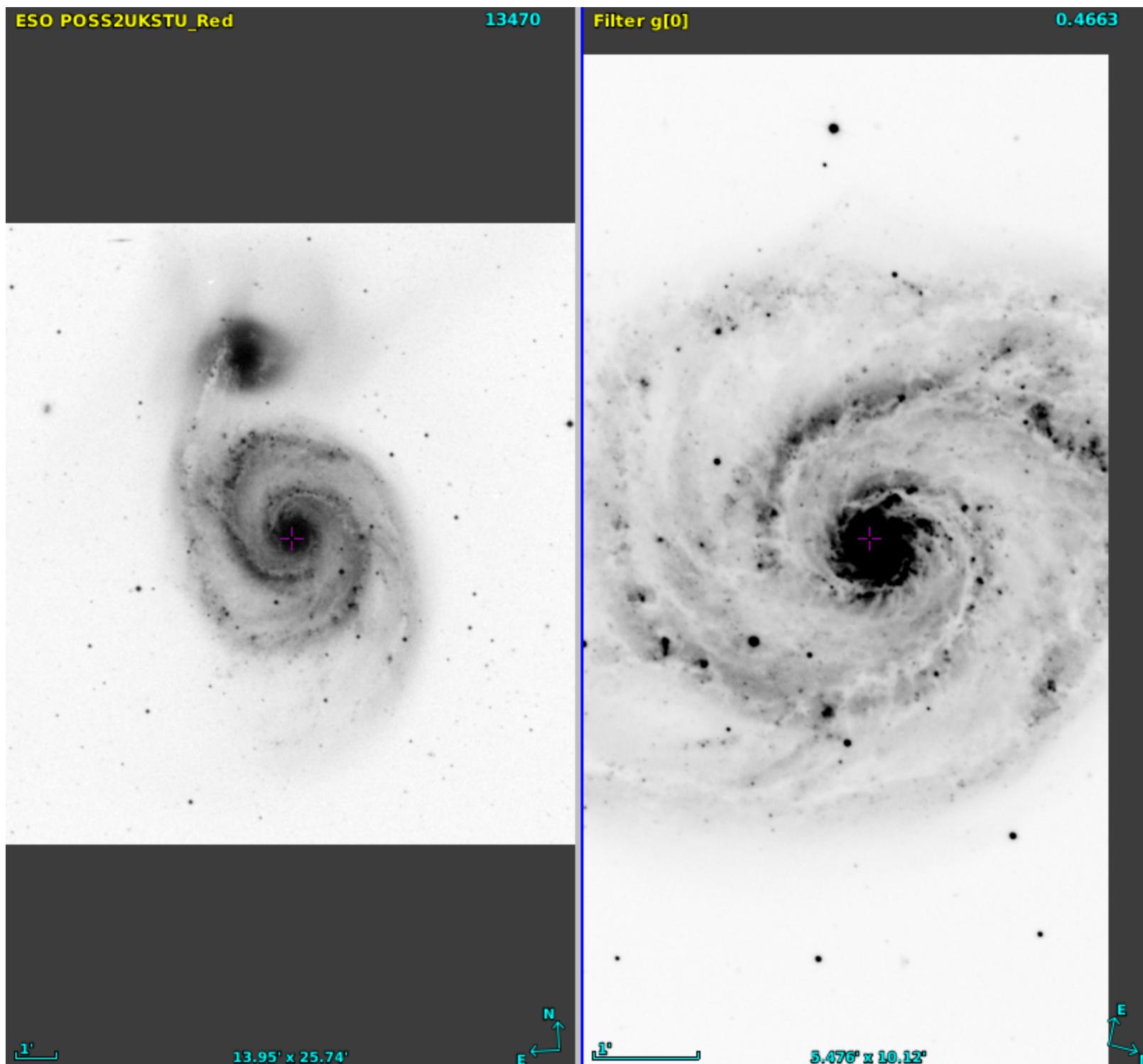
TESS



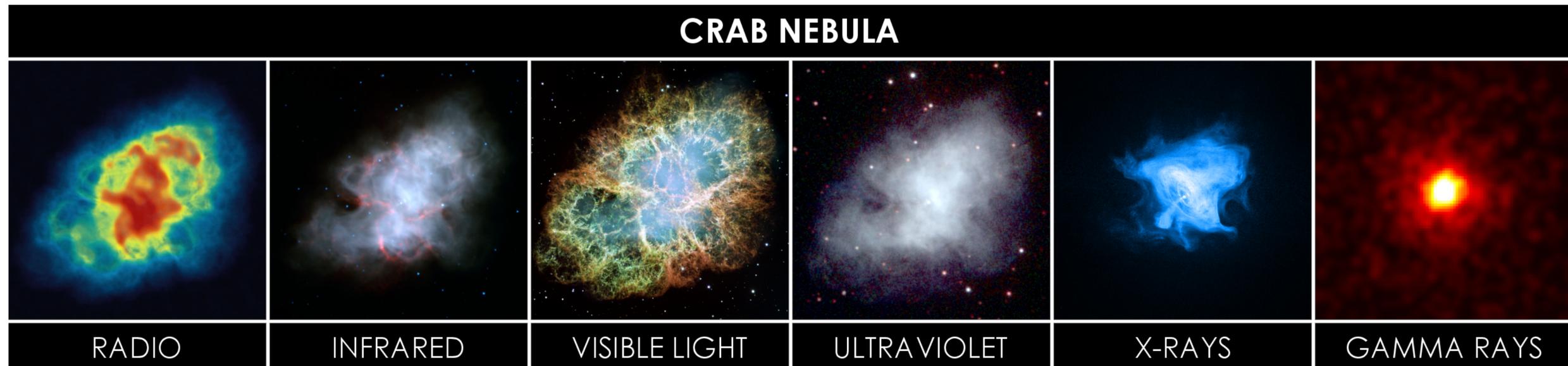
See All of MAST's Missions and Data

- Portal
- Catalogue Facility
- Community Forum

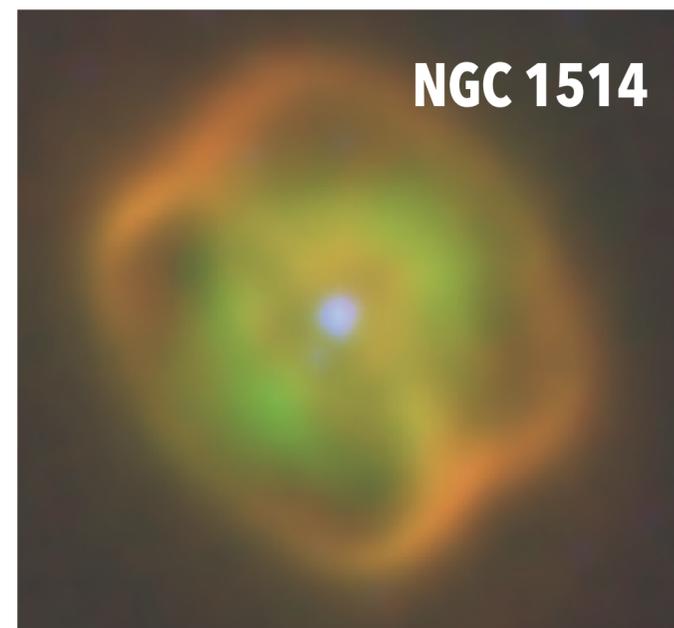
THE PROBLEM: INTEROPERABILITY



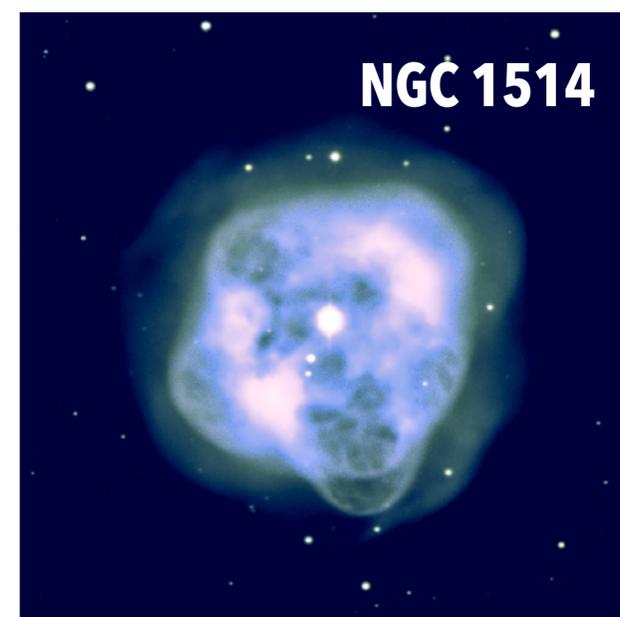
MULTI-WAVELENGTH ASTRONOMY



Credits: NRAO/AUI and M. Bietenholz; NRAO/AUI and J.M. Uson, T.J. Cornwell (radio); NASA/JPL-Caltech/R. Gehrz / University of Minnesota (infrared); NASA, ESA, J. Hester and A. Loll / Arizona State University (visible); NASA/Swift/E. Hoversten, PSU (ultraviolet); NASA/CXC/SAO/F.Seward et al. (X-rays); NASA/DOE/Fermi LAT/R. Buehler (gamma rays).

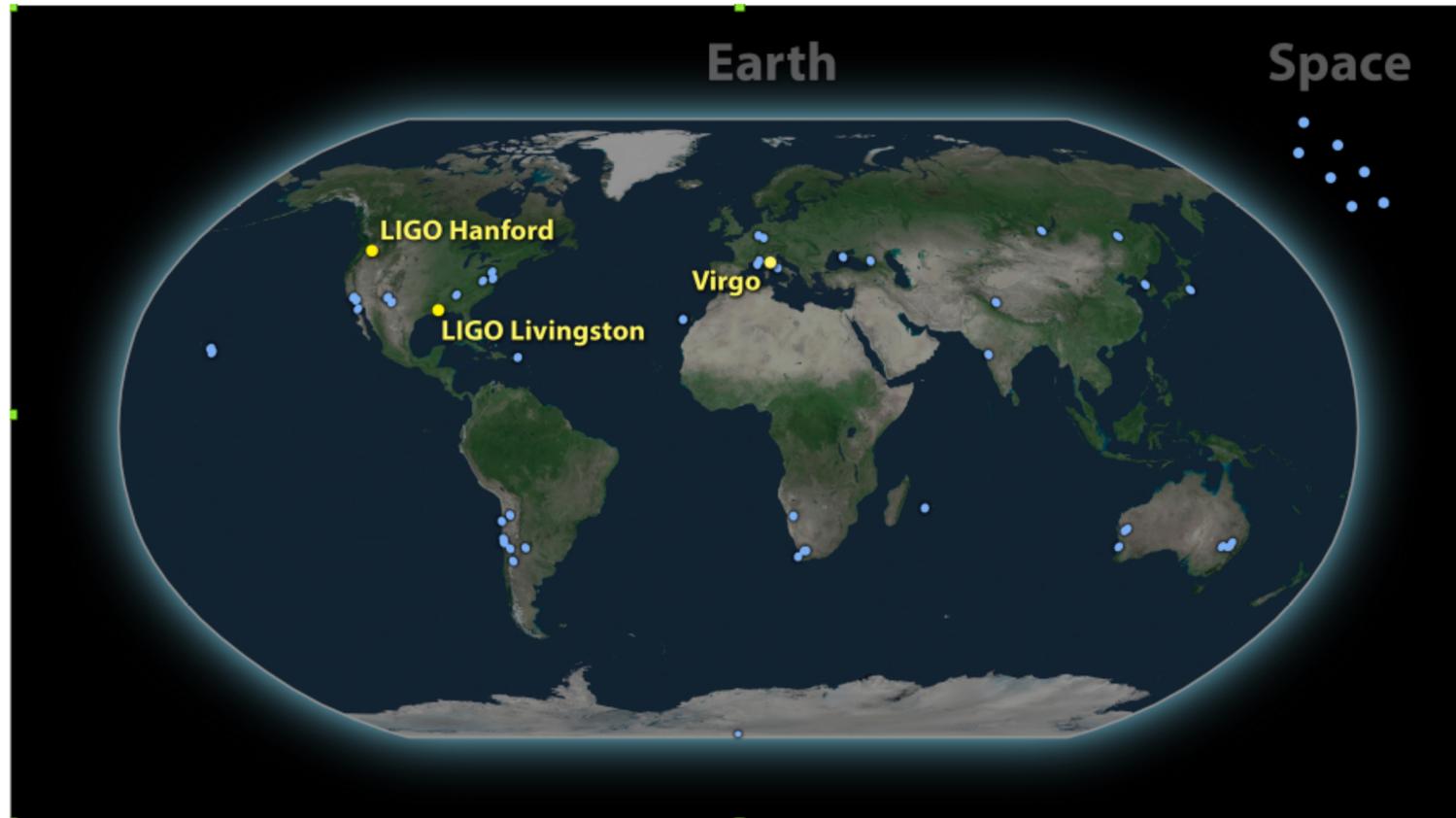


Credits: Ressler et al. 2010



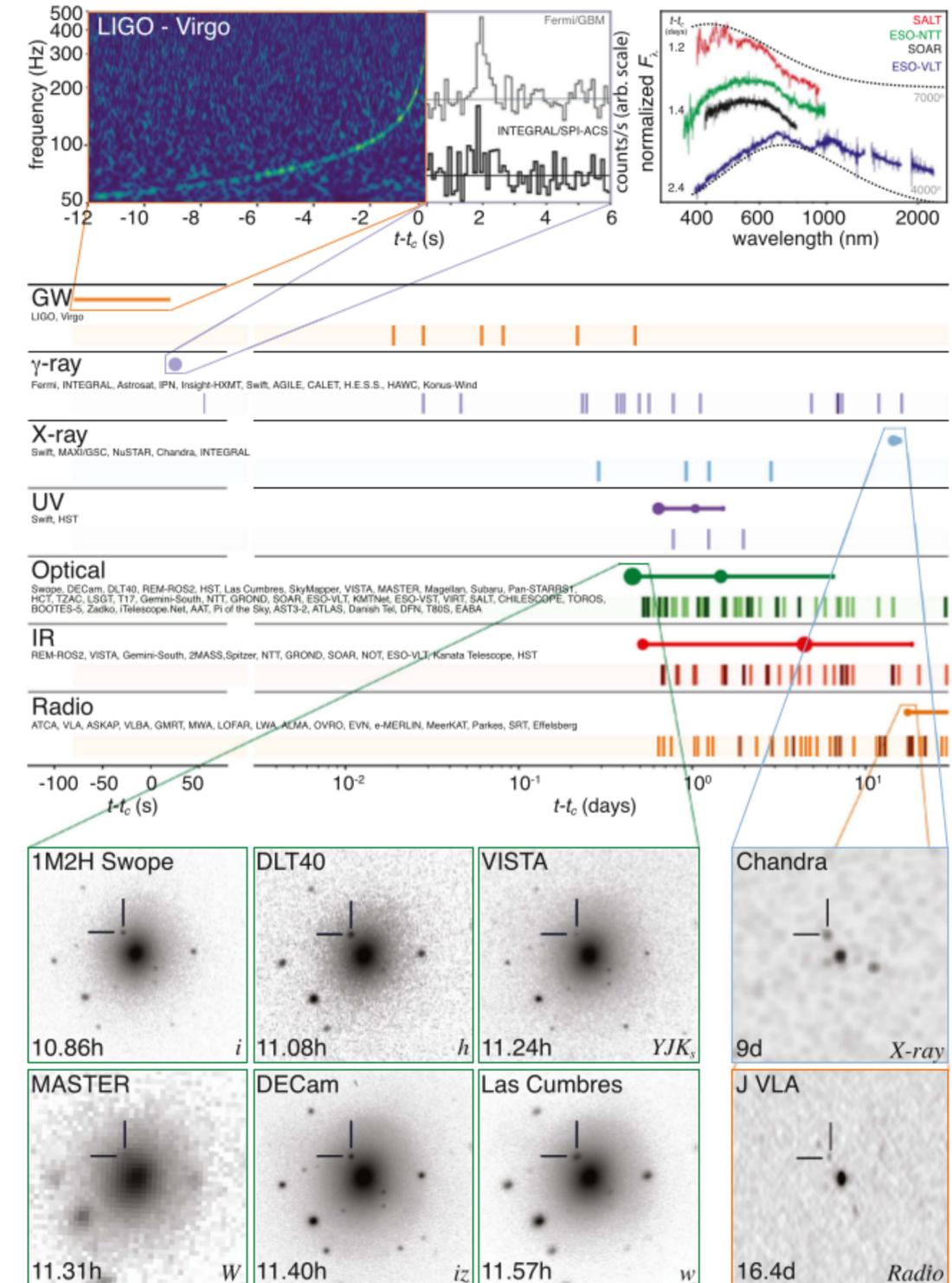
Credits: Jones et al. 2017

MULTI-WAVELENGTH ASTRONOMY



THE ASTROPHYSICAL JOURNAL LETTERS, 848:L12 (59pp), 2017 October 20

Abbott et al.

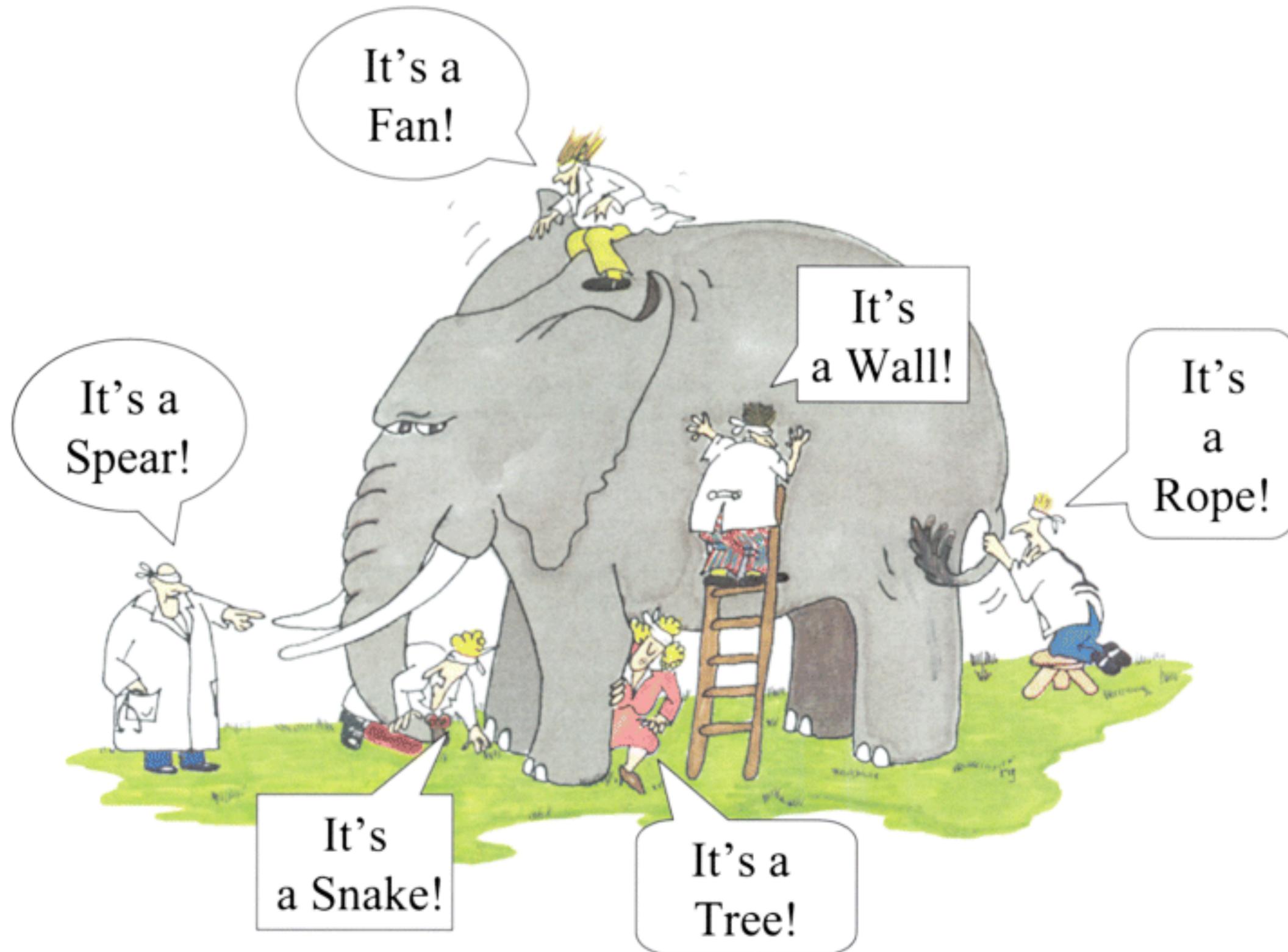


GW170817

~ 4000 astronomers

~ 900 groups

MULTI-WAVELENGTH ASTRONOMY



MULTI-WAVELENGTH ASTRONOMY

THE ASTROPHYSICAL JOURNAL, 644: L75–L78, 2006 June 10
© 2006. The American Astronomical Society. All rights reserved. Printed in U.S.A.

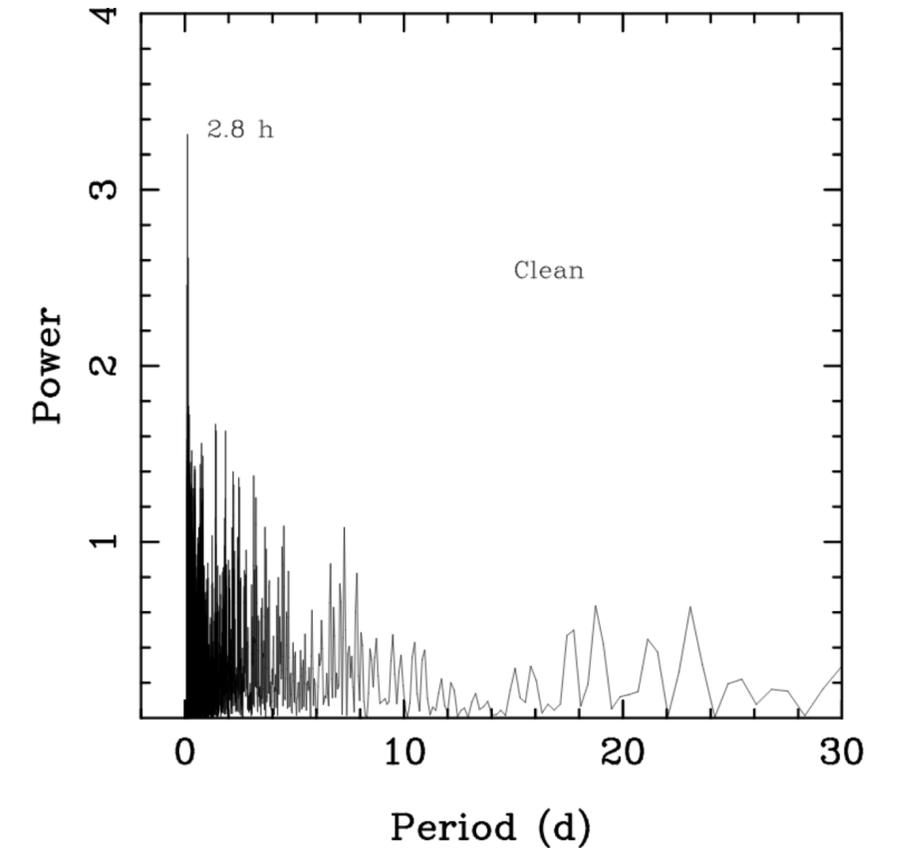
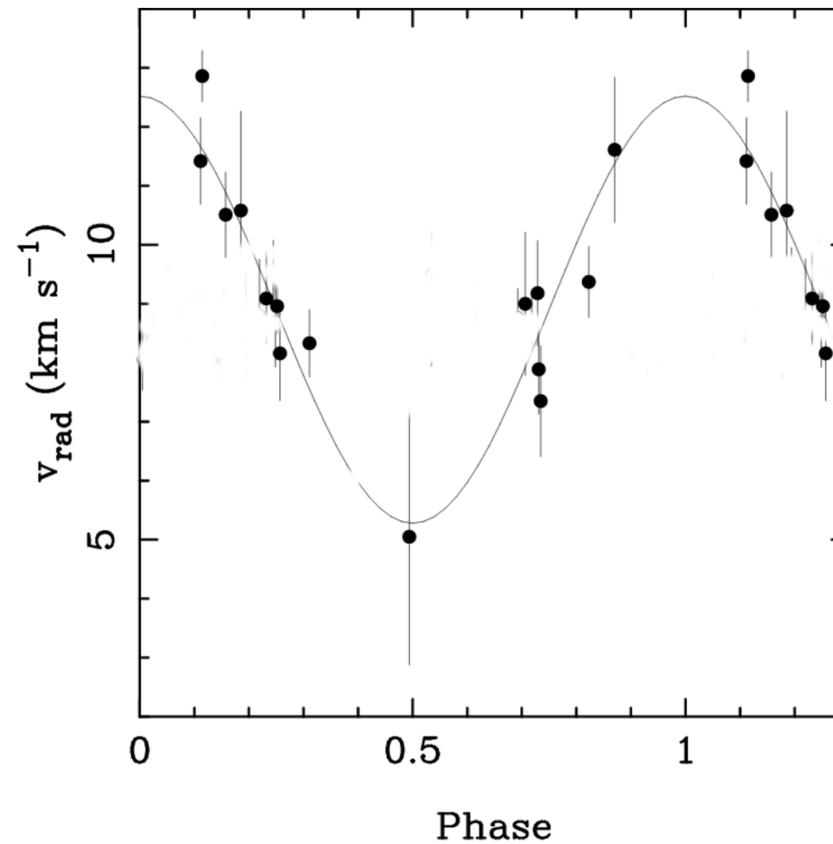
A MULTIWAVELENGTH RADIAL VELOCITY SEARCH FOR PLANETS AROUND THE BROWN DWARF LP 944-20

E. L. MARTÍN,^{1,2} E. GUENTHER,³ M. R. ZAPATERO OSORIO,⁴ H. BOUY,¹ AND R. WAINSCOAT⁵

Received 2006 April 10; accepted 2006 April 25; published 2006 May 26

Optical range (UVES/VLT):

- 14 nights /841 days
- Period 2.5 - 3.7h



MULTI-WAVELENGTH ASTRONOMY

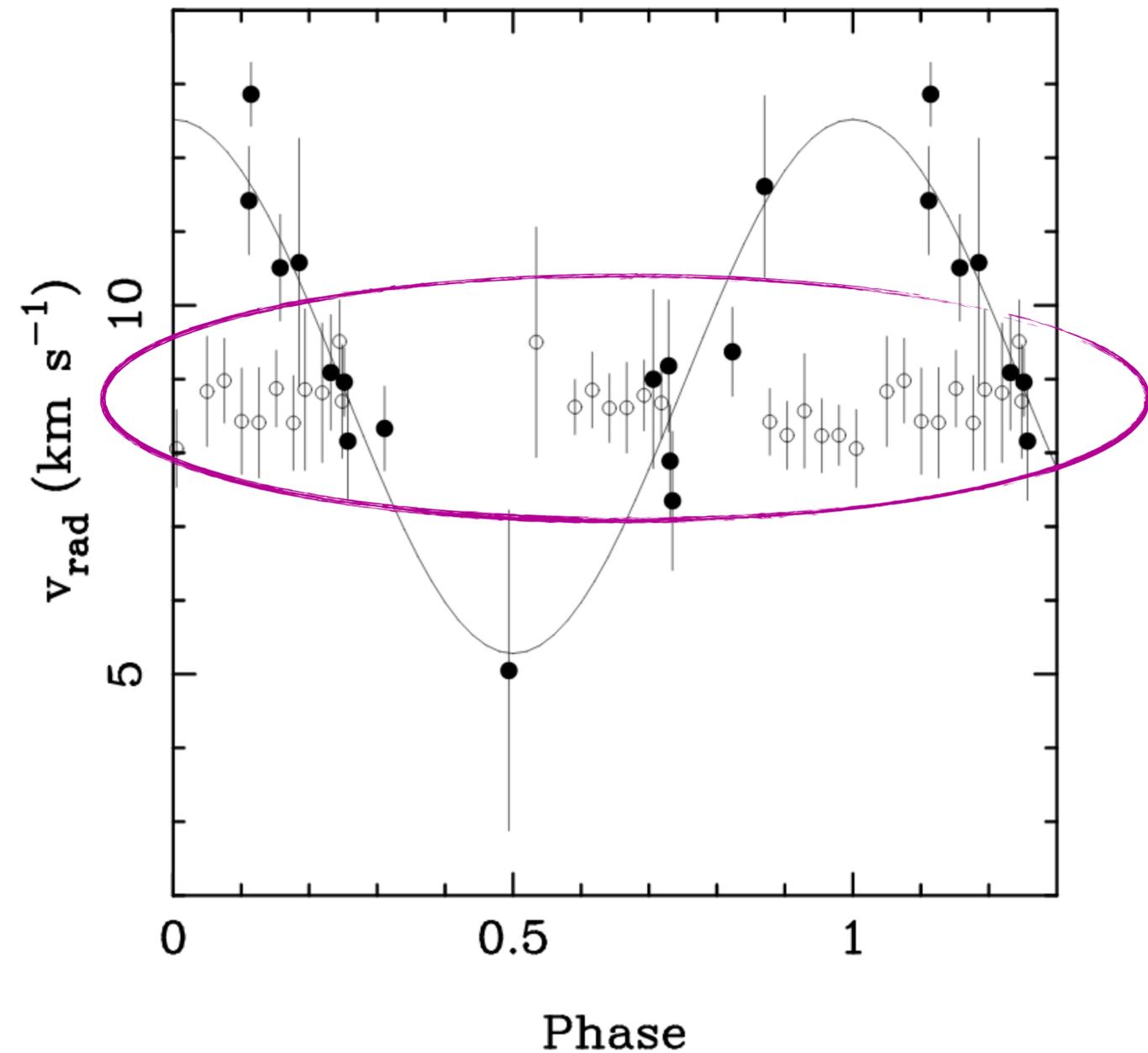
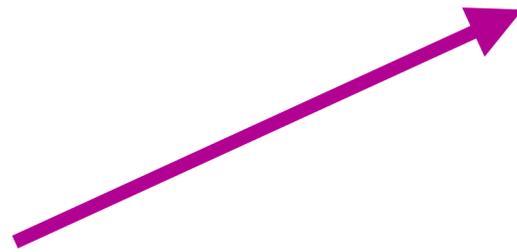
THE ASTROPHYSICAL JOURNAL, 644: L75–L78, 2006 June 10
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A MULTIWAVELENGTH RADIAL VELOCITY SEARCH FOR PLAN

E. L. MARTÍN,^{1,2} E. GUENTHER,³ M. R. ZAPATERO OSOR

Received 2006 April 10; accepted 2006 April 25;

Infrared range



THE GOAL: MAKING DATA **F.A.I.R**

F

A

I

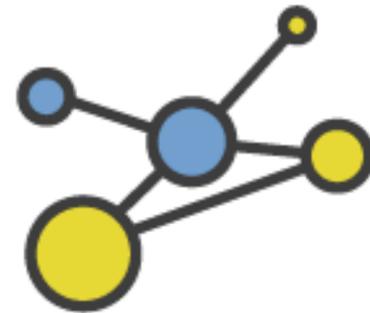
R



Findable



Accessible



Interoperable



Reusable

THE INTERNATIONAL VIRTUAL OBSERVATORY ALLIANCE

Virtual Observatories of the Future

Caltech campus, Pasadena, Calif., USA
June 13 - 16, 2000
<http://astro.caltech.edu/nvoconf>
Email inquiries: nvoconf@astro.caltech.edu

MPA/ ESO/ MPE Joint Astronomy Conference

MINING THE SKY

July 31 - August 4, 2000
Garching, Germany



THE VIRTUAL OBSERVATORY ROADMAP

Development
of standards

1

Uptake of standards
by data centres

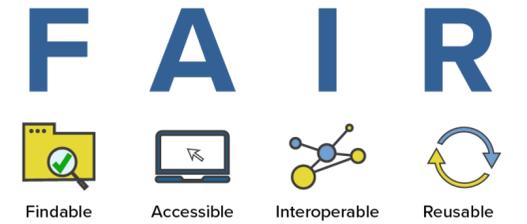
2

VO-tools

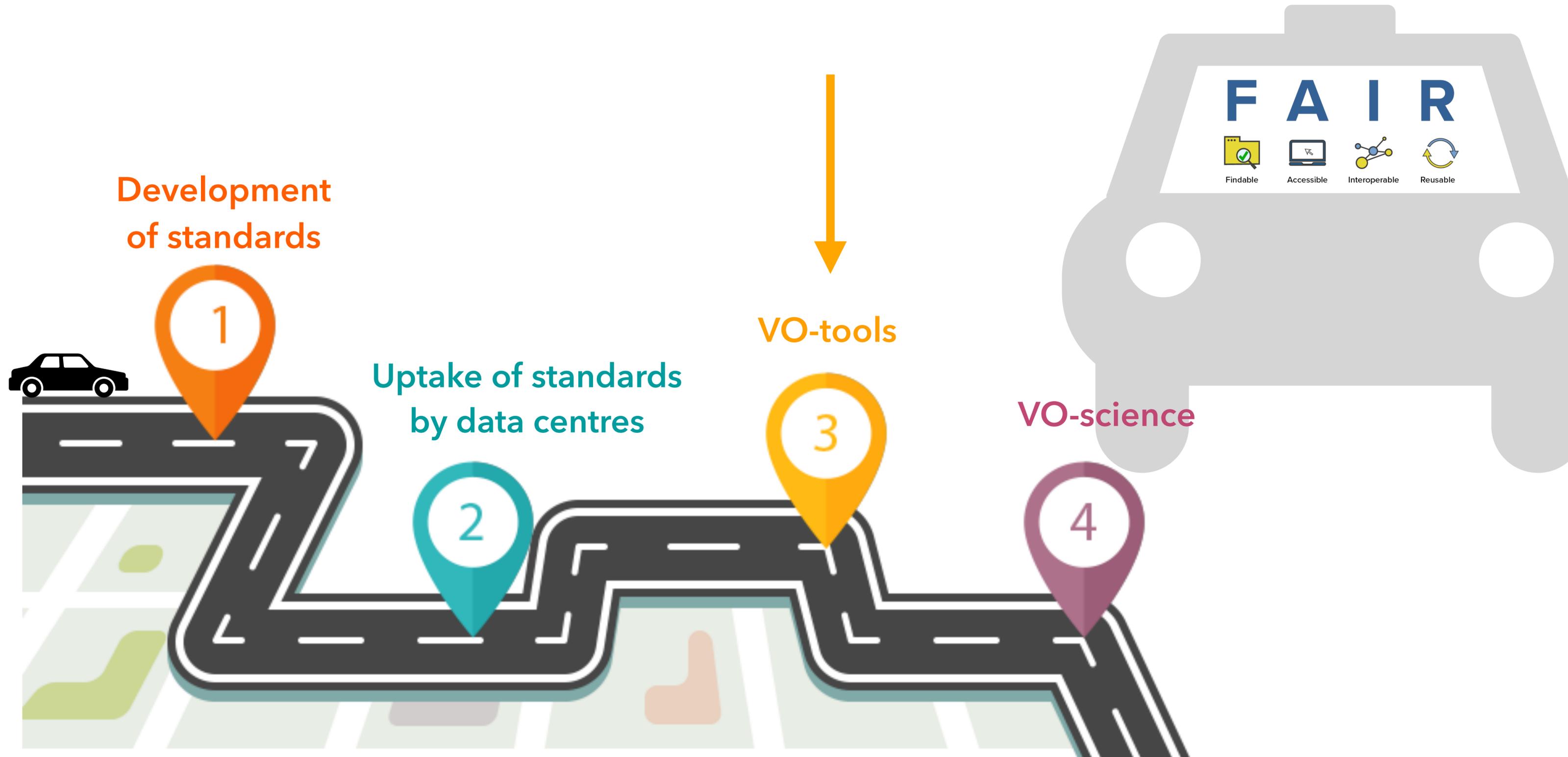
3

VO-science

4



THE VIRTUAL OBSERVATORY ROADMAP



THE POWER OF VO: **ALADIN**

A screenshot of the ALADIN v10.0 software interface. The main window displays a star field with a prominent nebula. The interface includes a menu bar (File, Edit, Image, Catalog, Overlay, Coverage, Tool, View, Interop, Help), a toolbar with various icons, and a data discovery tree on the left. The central panel shows a star field with a nebula, and the right panel shows a data discovery tree and a zoomed-in view of a star. The bottom status bar shows coordinates and other information.

Aladin v10.0

File Edit Image Catalog Overlay Coverage Tool View Interop Help

Available data → 24028 / 24029
● in view ● out view

Collections → 24028
Image → 477
Gamma-rav → 23
X → 23
UV → 27
Optical → 96
Infrared → 179
Radio → 67
Gas-lines → 45
Xrav → 1
X-rav → 16
Data base → 59
Catalog → 22530
VizieR → 21095
I-Astrometric Data → 267
II-Photometric Data → 346
AllWISE Data Release (Cutri+ 2
The Pan-STARRS release 1 (PS
2MASS All-Sky Catalog of Poin
The ISO GAL Point Source Catalo
Revised catalog of GALEX UV sou
KIDS-ESO-DR3 multi-band sour
AAVSO Photometric All Sky Sur
WISE All-Sky Data Release (Cu
GALEX-DR5 (GR5) sources from
2MASS 6X Point Source Workir
UKIDSS-DR9 LAS, GCS and DXS
AKARI/IRC mid-IR all-sky Surve
AKARI/FIS All-Sky Survey Point
SkyMapper Southern Sky Surve
The VISTA Hemisphere Survey
IRAS catalogue of Point Source
JMMC Stellar Diameters Catalog
IRAS PSC/FSC Combined Catal
IRAS Faint Source Catalog, [b]
GLIMPSE Source Catalog (I + I
The SDSS Photometric Catalog
TASS Mark IV patches photome
Catalog of Infrared Observations
IRAS Point Source Reject Catal
XMM-OM Serendipitous Source
General Catalogue of Photome
VPHAS+ DR2 survey (Drew+,
XMM-OM Serendipitous Source S
VLT Survey Telescope ATLAS (I
UBV Photoelectric Catalog, dat
UBV Photometry of O & B Star
Palomar Transient Factory (PTF
Homogeneous Means in the UB

Command
DSS SDSS 2MASS Simbad NED +
CDS/P/DSS2/color

Frame ICRS Projection Spheric

select
pan
dist
phot
draw
tag
moc
spect
filter
cross
xy
rgb
assoc
crop
cont
pixel
prop
del

Data discovery tree
of available collections (left panel)
● ingreen: the collections visible in the current view, in orange the outside collections
● Type un key word, a mission name, etc in the select field to keep only relevant collections
● Use the predefined filters to select more specifically the collections
● Create and store your own filters thanks to the right button
Activate the view mode to hide temporary the collections outside the view.

epoch -
size -
dens. -
opac. -
zoom -

Barnard 33
05:40:59.69 -02:16:12.0
1,306° x 53.07'

15'
1.306° x 53.07'

coll. view scan filter
grnd study wink north hdr multiview match

© 2017 Université de Strasbourg/CNRS - by CDS - Distributed under GNU GPL v3
0 sel / 0 src 34fps / 299Mb

THE POWER OF VO: ALADIN

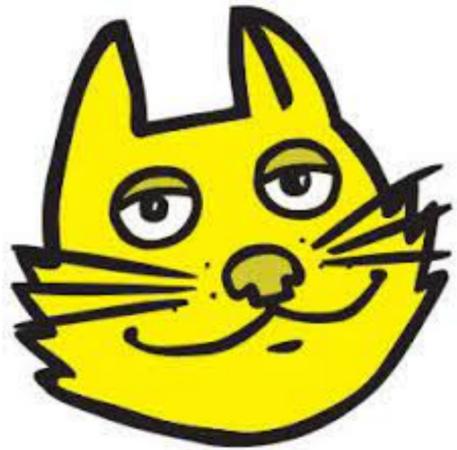


The screenshot displays the ALADIN v10.0 interface, a web-based visualization tool for astronomical data. The main window is divided into four panels, each showing a different color representation of the same astronomical field:

- Top-left panel:** CDS/P/DSS2/color. Shows a wide-field view with a bright star and a crosshair.
- Top-right panel:** CDS/P/PanSTARRS/DR1/color-z-zg-g. Shows a zoomed-in view of the star and surrounding field.
- Bottom-left panel:** xcatdb/P/XMM/PN/color. Shows a zoomed-in view of the star and surrounding field.
- Bottom-right panel:** CDS/P/2MASS/color. Shows a zoomed-in view of the star and surrounding field.

The interface includes a menu bar (File, Edit, Image, Catalog, Overlay, Coverage, Tool, View, Interop, Help), a command line, and a sidebar with a tree view of available data. The sidebar shows a list of data collections, including HST-wideV, GTC Public Archive, DECaLS, DECaPS, DES, HSC, IPHAS, MAMA, PanSTARRS, J-PLUS-DR1, MINIJ-PAS-PDR201912, BASS, DES DR1 LIneA color, Swift, UVOT, Infrared, Radio, Gas-lines, X-ray, INTEGRAL, and XMM. The bottom status bar shows the current view: [View B2] - CDS/P/2MASS/color, and the bottom right corner displays the current coordinates: 085.204211 -01.96207 ICRS.

THE POWER OF VO: **TOPCAT**



TOPCAT interface showing a table list, current table properties, and a plane plot.

Table List
4: TAP_4_gaiadr1.tgas_soi

Current Table Properties
Label: TAP_4_gaiadr1.tgas_source,gaiadr1.tmass_best_nei...
Location: TAP_4_gaiadr1.tgas_source,gaiadr1.tmass_best_neighbour,gaiadr1.tmass_...
Name: sync
Rows: 10,000
Columns: 3
Sort Order: ↑
Row Subset: All
Action: (no action) Broadcast Row

Plane Plot
g_mag_abs vs g_min_ks

Table Access Protocol (TAP) Query interface showing metadata and ADQL text.

Metadata

Name	Service	Schema	Table	Columns	FKeys	Hints
solution_id				BIGINT	✓	
designation				VARCHAR	✓	
source_id				BIGINT	✓	
random_index				BIGINT	✓	
ref_epoch				DOUBLE	✓	yr
ra				DOUBLE	✓	deg
ra_error				DOUBLE	✓	mas
dec				DOUBLE	✓	deg
dec_error				DOUBLE	✓	mas
parallax				DOUBLE	✓	mas
parallax_error				DOUBLE	✓	mas
parallax_over_error				REAL	✓	
pmra				DOUBLE	✓	mas.yr**-1
pmra_error				DOUBLE	✓	mas.yr**-1
pmdec				DOUBLE	✓	mas.yr**-1
pmdec_error				DOUBLE	✓	mas.yr**-1
ra_dec_corr				REAL	✓	
ra_parallax_corr				REAL	✓	
ra_pmra_corr				REAL	✓	

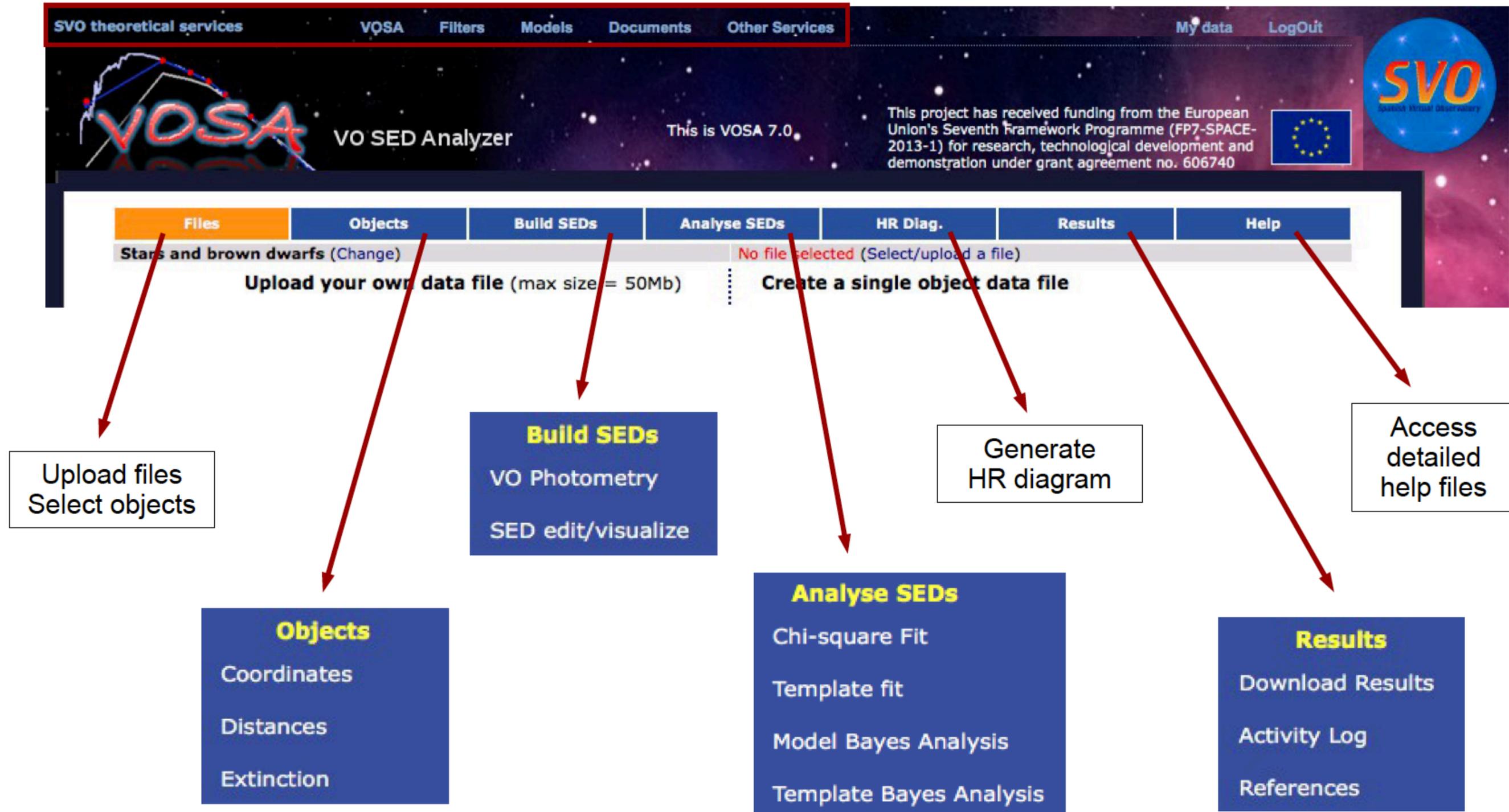
Service Capabilities
Query Language: ADQL-2.0 Max Rows: 300000 (default) Uploads: 100Mb

ADQL Text
Mode: Synchronous

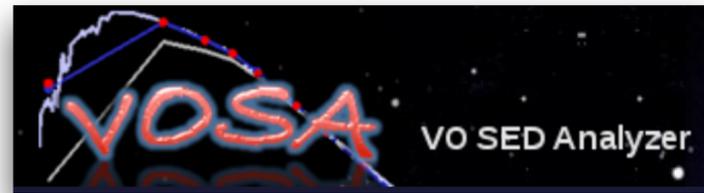
```
SELECT TOP 10000 gaia.source_id,  
gaia.phot_g_mean_mag + 5 * log10(gaia.parallax) - 10 AS g_mag_abs ,  
gaia.phot_g_mean_mag - tmass.ks_m AS g_min_ks  
FROM gaiadr1.tgas_source AS gaia  
INNER JOIN gaiadr1.tmass_best_neighbour AS xmatch  
ON gaia.source_id = xmatch.source_id  
INNER JOIN gaiadr1.tmass_original_valid AS tmass  
ON tmass.tmass_oid = xmatch.tmass_oid  
WHERE gaia.parallax/gaia.parallax_error >= 5 AND  
ph_qual = 'AAA' AND  
sqrt(power(2.5 / log(10) * gaia.phot_g_mean_flux_error  
/ gaia.phot_g_mean_flux, 2) ) <= 0.05 AND  
sqrt(power(2.5/log(10)*gaia.phot_g_mean_flux_error  
/ gaia.phot_g_mean_flux, 2)  
+ power(tmass.ks_msigcom, 2)) <= 0.05
```

Run Query

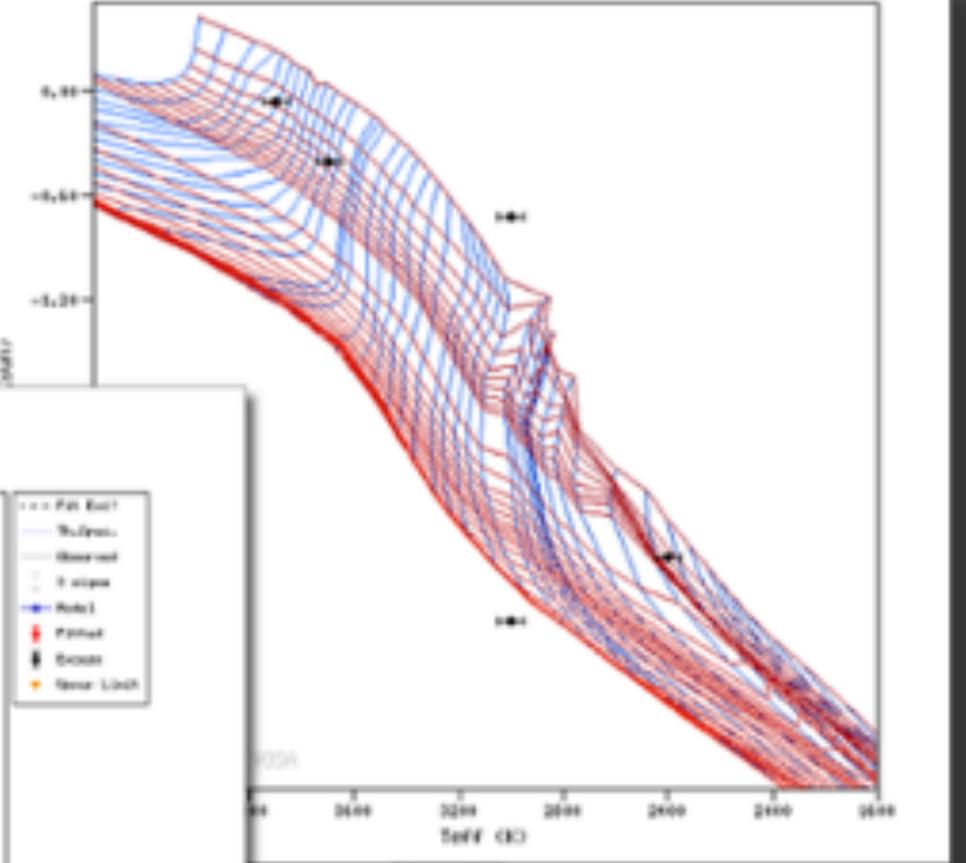
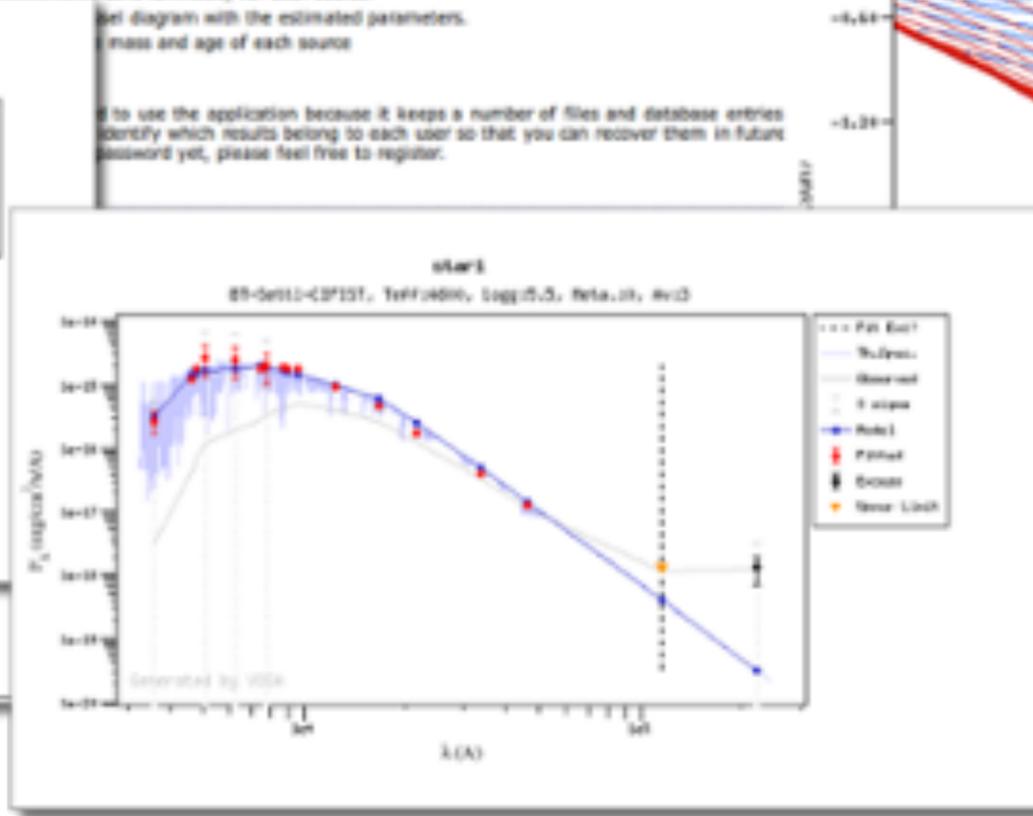
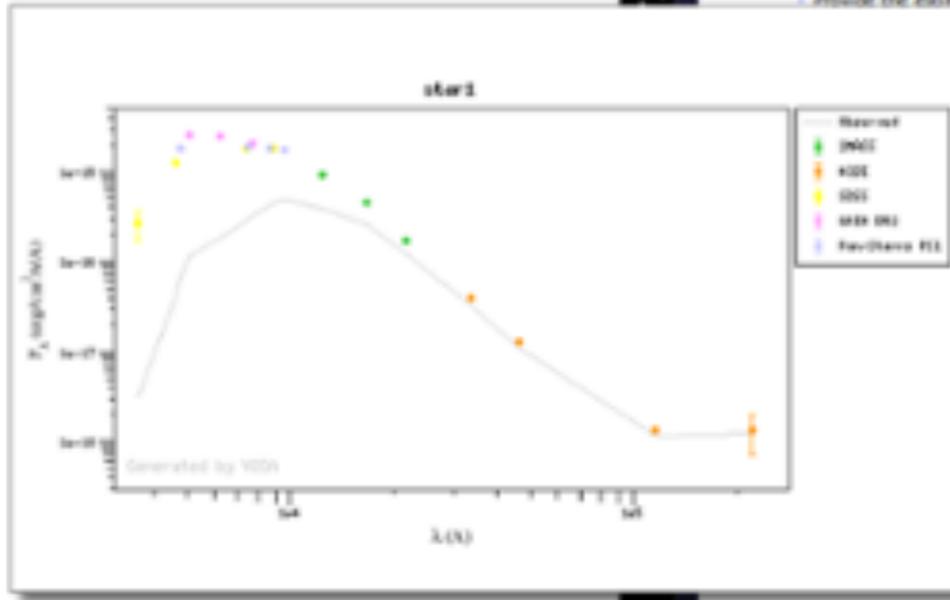
THE POWER OF VO: **VOSA**



THE POWER OF VO: **VOSA**

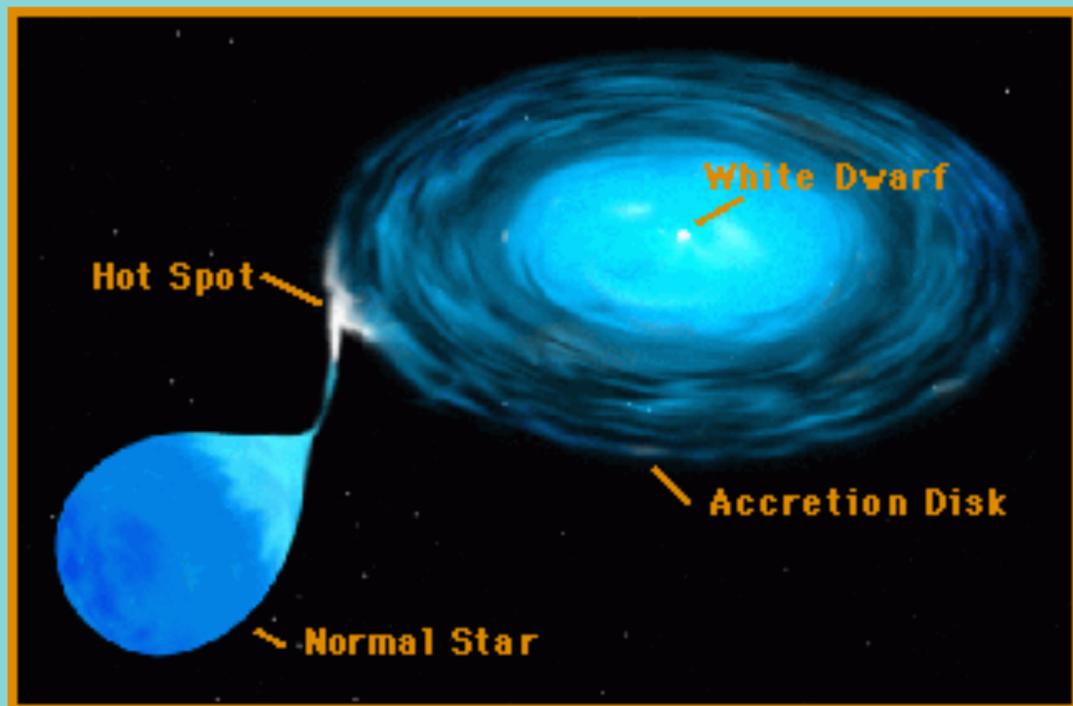


- VOSA (VO Sed Analyzer) is a tool designed to perform the following tasks in an automatic manner:
- Read user photometry tables.
 - Query several photometrical catalogs accessible through VO services (increases the wavelength coverage of λ to be analyzed).
 - Query VO-compliant theoretical models (spectra) and calculate their synthetic photometry.
 - Perform a statistical test to determine which model reproduces best the observed data.
 - Use the best-fit model as the source of a bolometric correction.
 - Provide the estimated bolometric luminosity for each source.



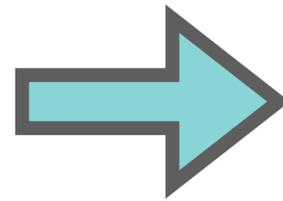
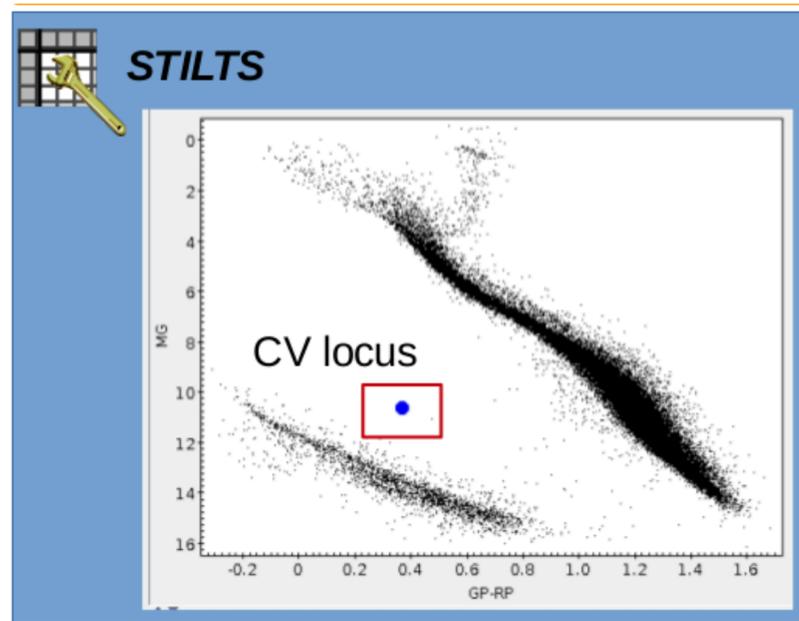
THE POWER OF VO: **TRANSIENTS**

WORKFLOW: Discovering new **Cataclysmic Variables** with VO tools



- Halpha emission due to accretion
- Close binaries (WD+Main seq.) → composite SEDs
- Well defined locus in the HR diagram

THE POWER OF VO: TRANSIENTS



VO Spectrum Services (VO: SSAP)

Services	Results
6	19

Service name	Results
DFBS SSAP	2
FUSE	1
LAMOST DR2 SSAP	6
LAMOST DR3 SSAP	6

Halpha emission

VOSA (VO: ConeSearch)

Koester WD models, Teff:8000, logg:9.5

Flux (erg/cm²/s/A)

Wavelength (A)

Generated by VOSA

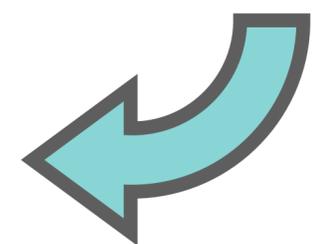
If excess
→ 2-body fitting

Koester, Teff:9500, logg:9.5 BT-Setti-CIFIST, Teff:2600, logg:5, Meta.:0 Av:0

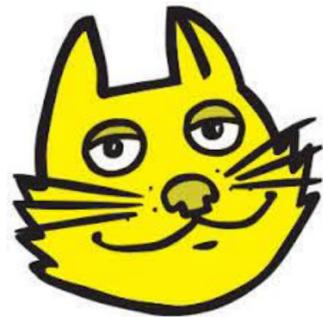
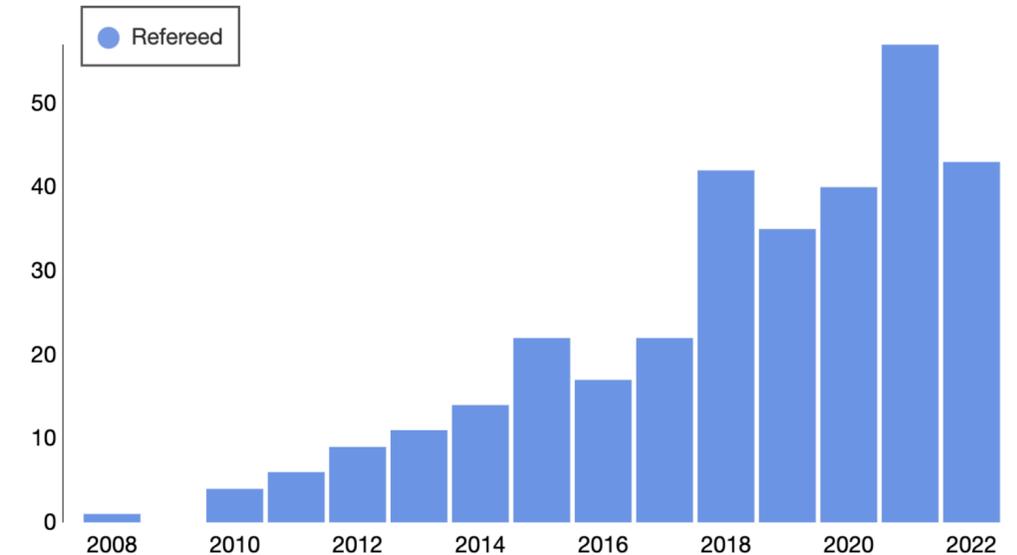
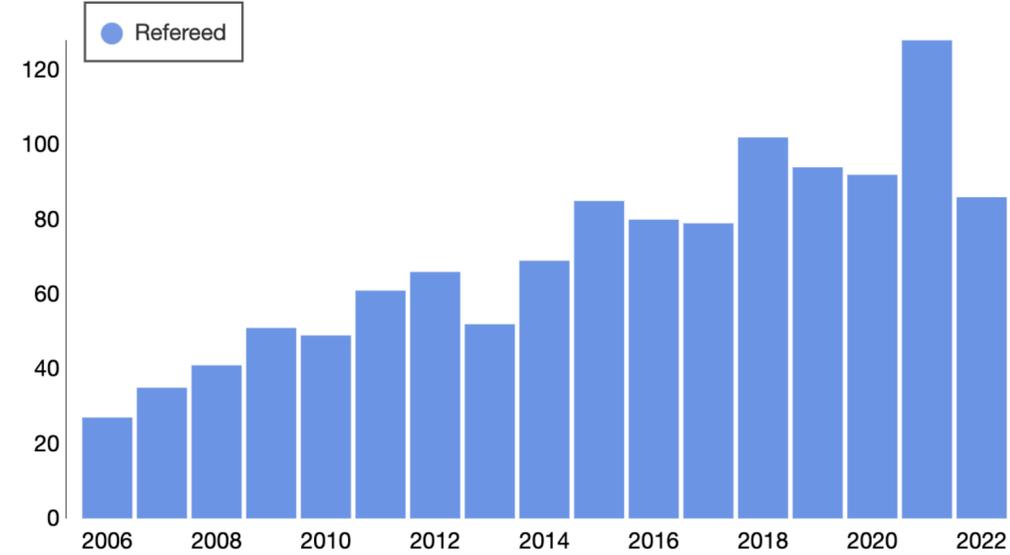
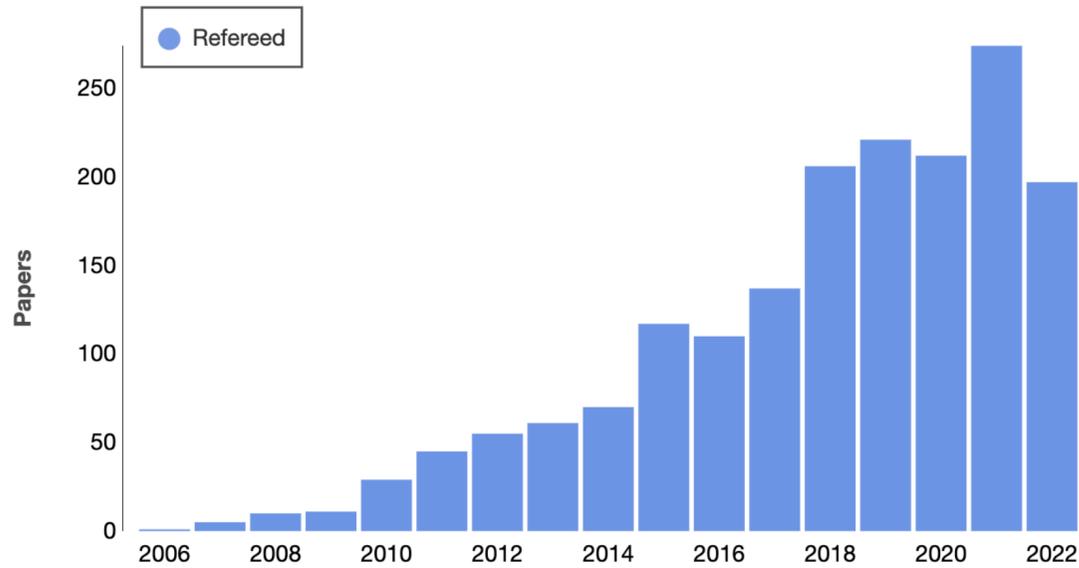
Flux (erg/cm²/s/A)

Wavelength (A)

Generated by VOSA



HOW USED ARE VO-TOOLS?



F.A.I.R. PRINCIPLES BEYOND ASTROPHYSICS



We agree and endorse the principles below when dealing with research results stemming from EU funded research grants related to COVID-19 on prevention (including vaccines), testing and treatment:

- 1** Make the generated results, whether tangible or intangible, **public and accessible** without delay, for instance on the Horizon Results Platform, on an existing IP sharing platform, or through an existing preprint pool.
- 2** Make scientific papers and research data available in **open access** without delay and following the **FAIR principles** via preprint servers or public repositories, with rights for others to build upon the publications and data and with access to the tools needed for their validation. In particular, make COVID-19 research data available through the European COVID-19 Data Platform.
- 3** Where possible, grant for a limited time³, **non-exclusive royalty free licences** on the intellectual property resulting from EU-funded research. These non-exclusive royalty free licenses shall be given in exchange for the licensees' commitment to rapidly and broadly distribute the resulting products and services under fair and reasonable conditions to prevent, diagnose, treat and contain COVID-19.

ISBN 978-92-76-20627-9, doi:10.2777/661122, KI+01-20-430-EN-N

The FAIR principles are fulfilled in Astronomy thanks to the Virtual Observatory making reality what in other disciplines is just a dream

HILL: I felt a little bit like I stepped into the future when I went to a meeting of the [virtual observatories](#) in astronomy. They've already done a lot of the things that we're aiming to do in the sense that they've got observatories from all around the world that are sharing, through a common infrastructure, the data that they're measuring. And they're integrating it into a unified picture of the sky and our universe. These same data contribute to building large-scale simulations of the universe. I think that's extremely inspiring. If we can have a common infrastructure that allows us to have a globally integrated view of the data being produced, and the tools to run large-scale simulations from the data, we will really have made progress in neuroscience.



Human Brain Project

Science ▾

Platforms ▾

Collaborate ▾

Follow HBP ▾

About ▾

Education & Training ▾

Welcome to the Human Brain Project

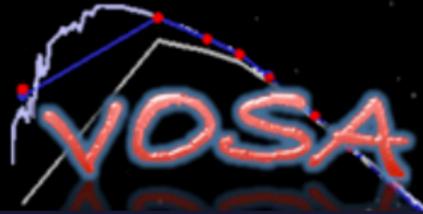
The Human Brain Project aims to put in place a cutting-edge research infrastructure that will allow scientific and industrial researchers to advance our knowledge in the fields of neuroscience, computing, and brain-related medicine

[Learn more about the project](#)

FINAL REMARKS

**VIRTUAL OBSERVATORY
MAKES YOUR LIFE
EASIER!! 😎**

If you have any question please, let us know: svo-support@cab.inta-csic.es



VO SED Analyzer

This is VOSA 7.5

This project has received funding from the European Union's Seventh Framework Programme (FP7-SPACE-2013-1) for research, technological development and demonstration under grant agreement no. 606740



Email:	<input type="text"/>
Pass:	<input type="password"/>
<input type="button" value="Login"/>	



If you are a new user, please, [register](#).
If you don't remember your password, [click here](#).

VOSA (VO Sed Analyzer) is a tool designed to perform the following tasks in an automatic manner:

- Read user photometry-tables.
- Query several photometrical catalogs accessible through VO services (increases the wavelength coverage of the data to be analyzed).
- Query VO-compliant theoretical models (spectra) and calculate their synthetic photometry.
- Perform a statistical test to determine which model reproduces best the observed data.
- Use the best-fit model as the source of a bolometric correction.
- Provide the estimated bolometric luminosity for each source.
- Generate a Hertzsprung-Russel diagram with the estimated parameters.
- Provide an estimation of the mass and age of each source

[\(Take a look to the VOSA Help\)](#)

You need a username and password to use the application because it keeps a number of files and database entries with your results and we need to be able to identify which results belong to each user so that you can recover them in future sessions. If you don't have a username and password yet, please feel free to register.

<http://svo2.cab.inta-csic.es/theory/vosa/>